

Electron-electron weak interaction in atoms and ions

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We discuss parity-nonconservation effects in two-electron atoms and ions, arising as the result of electron-electron neutral weak currents. Calculations are presented for the energy levels and transition probabilities for excited states of two-electron ions, and crossing of levels of opposite parity is observed in the region of charge $Z \sim 37$, which is the most suitable for observation of these effects. In the vicinity of this crossing the quantities characterizing the degree of parity nonconservation are of the order $\sim 10^{-8}$.

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The observation of neutral weak currents^[1] has led to intensive searches for electron-nuclear weak interactions in atoms.^[2-4] All methods of detection of weak interactions between an electron and a nucleus in an atom at low energies are based on the assumption of nonconservation of parity in neutral currents. The electron-electron weak interaction in atoms should have, apparently, the same order of magnitude as the electron-nuclear interaction. In observation of electron-electron weak interactions it is necessary to consider processes in which the electron-nuclear weak interaction is suppressed. Such processes are the emission of photons from atomic levels which have orbital angular momenta l greater than unity.^[5]

Because of the centrifugal barrier, the wave functions of the electrons go to zero as r^l as $r \rightarrow 0$, where r is the distance from the center of the nucleus. The weak interaction, which does not conserve parity, has zero range and is proportional to $\mathbf{a} \cdot \mathbf{p}$, where \mathbf{a} is some axial vector composed of the spins of the interacting particles and \mathbf{p} is the momentum operator. The product of the initial and final wave functions of the electrons goes to zero as $r^{l_1+l_2}$, where l_1 and l_2 are the orbital angular momenta at the beginning and end of the process. The momentum \mathbf{p} removes one power of r . Therefore after integration over the nuclear volume the matrix element of the weak-interaction potential turns out to be proportional to $(R/a)^{l_2+l_1-1}$, where R and a are the radii of the nucleus in the atom. For interaction with the nucleus the matrix elements between S and P states turn out to be nonvanishing as $R/a \rightarrow 0$. The matrix elements of the electron-electron interaction include integration over the entire volume of the atom and do not have parametric smallness, which depends on the magnitude of the orbital angular momenta. Therefore the matrix elements of the transitions $P \rightarrow P$, $D \rightarrow S$, and so forth for the electron-nuclear weak interaction are suppressed in comparison

with $S \rightarrow P$ transitions by at least a factor $R/a \sim 10^5$, while the matrix elements of electron-electron transitions have their previous value.^[1] This statement remains valid even for large nuclear charges Z , where relativistic effects are important and it is necessary to use Dirac Coulomb functions.

The weak interaction U_w , which is invariant with respect to time reversal and which does not conserve spatial parity, can be constructed only from the product of vector and axial currents.^[2-4] The electron-electron interaction contains only two independent relativistic invariants:

$$U_w = 2^{-1/2} G g (\gamma_\mu)_1 (\gamma_\mu \gamma_5)_2 + 2^{-1/2} G h (\sigma_{\mu\nu} q_\nu)_1 (\gamma_\mu \gamma_5)_2 + (1 \leftrightarrow 2), \quad (1)$$

where $G = 1.0 \times 10^{-5} m_p^{-2}$, m_p is the proton mass, g and h are certain constants, and the subscripts 1 and 2 refer to the first and second interacting electrons. The identity of the electrons is taken into account by antisymmetrization of the wave functions of the initial and final states.

In the nonrelativistic approximation interaction (1) takes the form

$$U_w = 2^{-1/2} G g (\sigma_1 - \sigma_2) (\mathbf{v}_1^+ - \mathbf{v}_2^+) + 2^{-1/2} G (g+h) [\sigma_1 \times \sigma_2] \times (\mathbf{v}_1^- - \mathbf{v}_2^-) + (1 \leftrightarrow 2), \quad (2)$$

where $\mathbf{v}_i^+ = (\mathbf{p}_i + \mathbf{p}'_i)/2m$; σ_i , \mathbf{p}_i , and \mathbf{p}'_i are the Pauli spin matrices and the initial and final momenta of the i -th electron. The second term of (1) contributes only to the second term of Eq. (2), while the first term of (1) contributes to both terms of Eq. (2). Therefore for simplicity we have carried out all calculations by setting $h = 0$, $g = 1$.

In coordinate space the interaction (1) takes the form

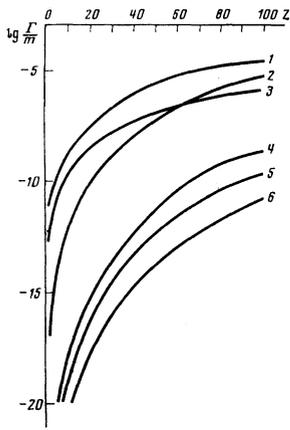


FIG. 3.

Thus, to obtain a large value of P it is necessary^[4] that the admixed level have a larger partial transition value than the main level ($\Gamma_{DP} \gg \Gamma_{PP}$). This leads to difficulties of observing the main transition line. In most experiments excitation of close-lying levels occurs with identical filling numbers. In view of the Breit-Wigner distribution of the radiation intensity as a function of frequency, the radiation from the admixed level will occur also at the frequency of the main transition. In view of the high intensity of radiation from the admixed level this can lead to an effective reduction of the value of P . In this case it is necessary to strive for a low population of the admixed level. This can be achieved by choosing as the object of study a level with long lifetime; by delaying the beginning of the measurement with respect to the moment of excitation of the ions, this permits an emptying of the admixed level with an insignificant change in the population of the investigated level.^[4] The lifetime of the 3^3P_2 level is roughly ten times that of the 3^3D_2 level, which determines our selection. Plots of the transition probability from the 3^3P_2 and 3^3D_2 levels to the low-lying levels are given in Fig. 3. These plots were calculated by us also by a completely relativistic method, as described in Ref. 9.

The probability of radiation for the transition $3^3P_2 \rightarrow 2^1P_1$ has the form

$$\Gamma_{PP}^W = \Gamma_{PP}(1 + Pn_s),$$

where Γ_{PP} is the transition probability in the absence of the weak interaction, P is defined by Eq. (4), n is the direction of photon emission, s is the photon spin, expressed in the well known way in terms of the polarization vector, $s = i(\mathbf{e} \times \mathbf{e}^*)$, or the initial spin of the ion. The weak-interaction matrix element, which determines P in accordance with Eq. (4), is calculated in the Appendix. For the strontium ion Sr the distance between the 3^3P_2 and 3^3D_2 levels is minimal and is given by $L = 6.5 \times 10^{-3} m \alpha^2$. The value of the parity-nonconservation coefficient (4) for strontium is $P = 1.8 \times 10^{-8}$. We have also calculated values of P for smaller Z values. In the region far from the point of crossing of the 3^3P_2 and 3^3D_2 levels, the main contribution to the second plot of Fig. 2b is from the intermediate level 3^1D_2 with the allowed E1 transition $3^1D_2 \rightarrow 2^1P_1$. In this case the ratio

of probabilities in Eq. (4) is $(\Gamma_{PP}/\Gamma_{DP})^{1/2} \sim (\alpha Z)^3$ and as a result the value P increases for small Z values to $P = 6.0 \times 10^{-8}$ at $Z = 2$ (the helium atom).

A plot of P as a function of the nuclear charge Z is given in Fig. 4. In the region of small Z , however, strong restrictions arise on the external electric fields, which as the result of Stark broadening of the lines lead to a reduction of P . In addition, the effect of collisions, which limits the density of the buffer gas,^[4] turns out to be important. For large Z these limitations do not arise.

Let us estimate the measurement time necessary for observation of the effect. To determine the degree of parity nonconservation P in a transition with partial probability B it is necessary to record $N \sim B^{-1}P^{-2}$ events. If we use an ion beam with density ρ traveling with velocity v in a volume V with a cross section S and length l , the number of decays N of the level in time t in this volume is

$$N = \rho v S t [1 - \exp(-l/v\tau_p)],$$

where τ_p is the total lifetime of the 3^3P_2 level. As a result we obtain the following expression for the time of observation t :

$$t = \tau/V\rho B P^2,$$

where

$$\tau = \frac{l}{v} \left[1 - \exp\left(-\frac{l}{v\tau_p}\right) \right]^{-1} = \begin{cases} l/v & (\tau_p \ll l/v), \\ \tau_p & (\tau_p \gg l/v). \end{cases}$$

In the case of experiments with ion beams^[10] we have $v \sim 10^8$ cm/sec, $\rho \sim 10^5$ cm⁻³, and for our transition we find $\tau \sim l/v$. If we assume $B \sim 10^{-5}$ this gives $t \sim (10^7/V [\text{cm}^3])$ sec. Hence it follows that to obtain an acceptable observation time $t \sim 10^3$ sec it is necessary to use a volume $V \sim 10^4$ cm³.

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APPENDIX

We have calculated the matrix elements of the operator U_W on the wave functions $\psi_{JMjj'11'}$ constructed according to jj coupling from the one-electron Dirac functions ψ_{njm} . To take into account the relativistic mixing

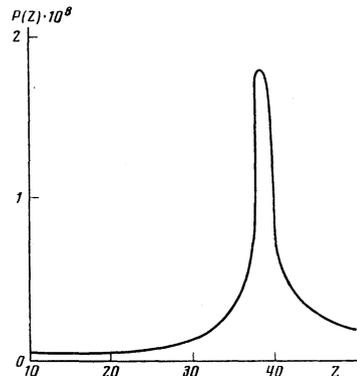


FIG. 4.

of the 3^3P_2 and 3^3D_2 states we have used wave functions of a more general form

$$\psi_{JM}(12) = \sum_{j'j''} \sum_{l'l''} a_j(jj'j'') \psi_{JMj'j''l'l''}, \quad (\text{A. 1})$$

where the mixing coefficient a_j are determined from the condition of minimum energy. The matrix element of most general form for the operator U_w on the functions $\psi_{JMj'j''l'l''}$ appears as follows:

$$\begin{aligned} & \langle J'M'j_1j_2'l_1l_2' | U_w | JMj_1j_2'l_1l_2' \rangle \\ &= -\frac{i}{\sqrt{2}} G(1-\mathcal{P}_{22'}) (1+\mathcal{P}_1\mathcal{P}_2-\mathcal{P}_1Q_2+Q_2\mathcal{P}_2'+\mathcal{P}_1\mathcal{P}_2'-\mathcal{P}_1\mathcal{P}_1' \\ & \quad -\mathcal{P}_1\mathcal{P}_1'Q_2Q_2') R(g_1g_1'f_2g_2') K(j_1j_2j_1'j_2' | l_1l_2l_1'l_2') \delta_{J,J'} \delta_{M,M'} \\ & \quad R(g_1g_1'f_2g_2') = \int_0^{\infty} g_{n_1n_1'}(r) g_{n_2n_2'}(r) f_{n_1n_1'}(r) g_{n_2n_2'}(r) r^2 dr, \\ & \quad \times K(j_1j_2j_1'j_2' | l_1l_2l_1'l_2') = (-1)^{l_1+l_2+l_1'+l_2'} \Pi_{j_1n_1l_1} \Pi_{j_2n_2l_2} \sum_l (2l+1) \\ & \quad \cdot \begin{pmatrix} l_1 & l_2 & l \\ 0 & 0 & 0 \end{pmatrix} \begin{pmatrix} l_1' & l_2' & l \\ 0 & 0 & 0 \end{pmatrix} \left\{ \begin{matrix} l_1 & l_2 & l \\ j_2 & j_1 & 1/2 \end{matrix} \right\} \left\{ \begin{matrix} l_1' & l_2' & l \\ j_2' & j_1' & 1/2 \end{matrix} \right\} \left\{ \begin{matrix} j_1 & j_2 & l \\ j_2' & j_1' & J \end{matrix} \right\} \\ & \quad + \left\{ \begin{matrix} l_2' & l_2 & j_2 & 1/2 \\ 1/2 & l & J & 1 & l_1' \end{matrix} \right\} \cdot \quad (\text{A. 2}) \end{aligned}$$

Here we have designated by g and f respectively the upper and lower components of the Dirac bispinors; the operators $\mathcal{P}_1, \mathcal{P}_1, \dots$ denote the substitution $g_1 \rightarrow f_1, g_1' \rightarrow f_1'$, and so forth in the integral $R(g_1g_1'f_2g_2')$ with simultaneous replacement of l_1, l_1', \dots by $\bar{l}_1 \equiv 2j_1 - l_1, \bar{l}_1' \equiv 2j_1' - l_1'$, etc. in the expression for K . The operator Q_2 signifies the inverse substitution $f_2 \rightarrow g_2, \bar{l}_2 \rightarrow l_2$. The symbol $\mathcal{P}_{22'}$ represents the operator of the permutation $n_2j_2l_2 \rightarrow n_2'j_2'l_2'$. In addition, we have used standard designations for the $3j, 6j,$ and $15j$ symbols. Finally,

$$\Pi_{l_1l_2l_1'l_2'} = [(2l_1+1)(2l_2+1)(2l_1'+1)(2l_2'+1)]^{1/2}.$$

Equation (A. 2) can be used to calculate the matrix element of the electron-electron weak interaction between two arbitrary states of two-electron ions with arbitrary nuclear charge Z .

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