CALCULATION OF THE $2p \rightarrow 1s$ TRANSITION ENERGY IN μ -MESIC ATOMS

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The $2p \rightarrow 1s$ transition energy in μ -mesic atoms is calculated for $13 \le Z \le 57$ and various models of charge distribution in the nucleus. When a Fermi-type model is employed in conjunction with the parameters found by Hofstadter et al from electron scattering on nuclei, the calculations are found to be in satisfactory agreement with the experimental data on the radiation from μ -mesic atoms with the indicated values of Z.

In connection with the recent experimental data on μ -mesic atoms, ^[1,2] we carried out a more accurate calculation of the $2p \rightarrow 1s$ transition energy in μ -mesic atoms for a number of elements between ₁₃Al and ₅₇La by means of a method developed by us earlier. ^[3,4] The calculations were performed for a smoothed homogeneous charge distribution (Fermi model):

$$\rho(r) = \rho_0 \{1 + \exp[(r - c)/z]\}^{-1}.$$
 (1)

The parameters of this distribution (its half-width c and surface layer thickness $t = 2z \ln 9 \approx 4.4 z$) were found by Hofstadter and co-workers from highenergy electron scattering on nuclei (see, for example, ^[5]). For the calculation, we took the average value of the parameters:

$$c = 1.08 A^{1/3}$$
 F and $t = 2.4$ F, (2)

where A is the atomic weight of the natural iso-tope mixture.

In the final results, we took into account the corrections for the energy levels as a result of the vacuum polarization, with allowance for the finite size of the nucleus.^[6] However, we did not take into account corrections for the polarization

of the nucleus by the meson. An estimate made by means of the Steinwedel-Jensen model^[7] shows that, with the accuracy of our calculations (somewhat greater than 0.1%), these corrections can be omitted. On the other hand, according to other earlier estimates of these quantities, these corrections can amount to 0.5-1% of the energy level for heavy elements.^[8] Their exact calculation requires individual consideration not only of each element but also of each isotope, and it is therefore very difficult.

The table lists the calculated energies (in keV) of the $1s_{1/2}(E_{1S_{1/2}})$ level; for the 2p level, the position of the center of gravity $(E_{2pc.g.})$ and the value of the relativistic splitting $(E_{2p_{3/2}} - E_{2p_{1/2}})$ are given.

The figure shows the relative deviation (in percent) of the experimental values of the E_e energy transition from the calculated values $E_t = E_{2pc.g.} - E_{1S_{1/2}}$ (black circles—data from ^[1], open circles—data from ^[2]). Shown in the same figure are the relative changes in E_t for changes in the individual parameters of the distribution (c by $\pm 2\%$ and t by $\pm 10\%$). Moreover, we also show the

Element	$-E_{1S_{1/2}}$	- <i>E</i> ₂ <i>p</i> c.g.	$E_{1p_{3/2}} - E_{2p_{1/2}}$	Element	$-E_{1s_{1/2}}$	- <i>E</i> ₂ <i>p</i> c.g.	$E_{2p_{3/2}} - E_{2p_{1/2}}$
13A1 16S 10K 20Ca 22Ti 24Cr 25Mn 26Fe 27Co 28Ni 29Cu 30Zn	$\begin{array}{r} 465\\ 695\\ 966\\ 1065\\ 1271\\ 1493\\ 1609\\ 1730\\ 1851\\ 1979\\ 2104\\ 2236\end{array}$	$\begin{array}{c} 119\\ 180\\ 255\\ 282\\ 342\\ 407\\ 442\\ 479\\ 516\\ 555\\ 596\\ 638\\ \end{array}$	$\begin{array}{c} 0.3 \\ 0.6 \\ 1.2 \\ 1.5 \\ 2.2 \\ 2.9 \\ 3.5 \\ 4.0 \\ 4.6 \\ 5.3 \\ 6.1 \\ 6.9 \end{array}$	³³ As ³⁷ Rb ⁴² Mo ⁴⁵ Rh ⁴⁶ Pd ⁴⁶ Pd ⁴⁷ Ag ⁴⁸ Cd ⁴⁹ In ⁵⁰ Sn ⁵¹ Sb ⁵⁶ Ba ⁵⁷ La	$\begin{array}{c} 2639\\ 3209\\ 3970\\ 4443\\ 4603\\ 4764\\ 4919\\ 5084\\ 5241\\ 5404\\ 6228\\ 6401 \end{array}$	$\begin{array}{r} 773\\973\\1256\\1443\\1508\\1574\\1642\\1711\\1782\\1854\\2235\\2314\end{array}$	$\begin{array}{c} 9.8\\ 14.8\\ 23.5\\ 29.8\\ 32.4\\ 35.1\\ 37.8\\ 40.5\\ 43.4\\ 45.5\\ 62.2\\ 65.8\end{array}$

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difference between the values of the energy of this transition calculated for a homogeneous distribution of nuclear charge, whose radius is R = $1.2 \text{ A}^{1/3}$ F, and values for the distribution (1) with the parameters (2).

For some elements, the values of the parameters c and t were found directly from experiments on electron scattering.^[5] However, the results of the calculation with these values of the parameters do not differ essentially from the results of the calculation with the averaged parameters (2), since the deviations of c and t from the mean values in these cases practically offset each other.

As is seen from the figure, the deviations of the experimental data from the calculated values are less than 1% in all cases (except for K, and in this case only for the data from ^[1]). Since the parameters c and t were determined by Hof-stadter with some uncertainty (about $\pm 2\%$ for c and $\pm 10\%$ for t), it can be concluded that the experiments on electron scattering and on the study of μ -mesic atoms are in good agreement with one another. Nevertheless, in the region between Ca and Zn, the nuclei are apparently more compact than should follow from the distribution (1) with the parameters (2).

We also attempted to calculate the charge distribution of the nucleus suggested by Ford and Hill (the so-called family II)^[9]:

$$\rho(r) = \rho_0 \left[1 - \frac{1}{2} \exp(-n) \right]^{-1} \\ \times \begin{cases} 1 - \frac{1}{2} \exp[-n(1 - r/c)], & r \leq c \\ \frac{1}{2} \exp[n(1 - r/c)], & r \geq c \end{cases}$$
(3)

Here c is the half-width of the distribution, while the thickness of the surface layer t is determined by the expression $nt = 2c \ln 5$. It turned out that, for the distribution (3) with the parameters (2), the transition energy agrees with the values obtained for the distribution (1) with the same value of c, but for t increased by 10%. As can be seen from the figure, the experimental data in this case is in poorer agreement with the distribution (3) than with the distribution (1). If it is concluded that the distribution (3) is in agreement with experiments on electron scattering at somewhat higher values of t (on the average, about 2.6 F), then the dis-agreement with experiment can only increase.

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